Velocity correlations in dense granular shear flows: Effects on energy dissipation and normal stress

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We study the effect of precollisional velocity correlations on granular shear flow by molecular dynamics simulations of an inelastic hard sphere system. Comparison of the simulations with kinetic theory reveals that the theory overestimates both the energy dissipation rate and the normal stress in the dense flow region. We find that the relative normal velocity of colliding particles is smaller than that expected from random collisions, and the discrepancies in the dissipation and the normal stress can be adjusted by introducing the idea of the collisional temperature, from which we conclude that the velocity correlation neglected in the kinetic theory is responsible for the discrepancies. Our analysis of the distributions of the precollisional velocity suggests that the correlation grows through multiple inelastic collisions during the time scale of the inverse of the shear rate. As for the shear stress, the discrepancy is also found in the dense region, but it depends strongly on the particle inelasticity.

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I. INTRODUCTION

Granular media can flow like fluid under certain situations. In the case of rapid granular flow, where the density is relatively low and interactions are dominated by instantaneous collisions, the kinetic theory of dense gases $[1]$ $[1]$ $[1]$ is extended to inelastic hard spheres to derive the constitutive relations $\lceil 2 \rceil$ $\lceil 2 \rceil$ $\lceil 2 \rceil$. In the theory, the density correlations is taken into account to some extent but not the velocity correlations in most cases. As the flow gets denser, however, the molecular chaos assumption becomes questionable. In addition, the interactions may no longer be approximated by instantaneous collisions, but enduring contacts take place around the random close packing fraction. A comprehensive granular rheology including the rather complicated dense regime has not been established yet.

During the last several years, careful experiments and large-scale molecular dynamics simulations have been done on dense granular flows $[3-6]$ $[3-6]$ $[3-6]$. One of the important model systems that has been intensively studied is the steady flow down a slope under the gravity, where we can control the ratio of the shear stress *S* to the normal stress *N* by changing the inclination angle θ . In this system, it has been found that the packing fraction ν in the bulk of the flow is constant and is determined solely by the inclination angle θ ; in other words, ν is independent of the total flow hight H and/or the roughness of the slope $[5,6]$ $[5,6]$ $[5,6]$ $[5,6]$.

This interesting feature has been qualitatively understood by using the Bagnold scaling $[7]$ $[7]$ $[7]$, which states that the shear stress *S* is proportional to the square of the shear rate $\dot{\gamma}$:

$$
S = m\sigma^{-1}A(\nu)\dot{\gamma}^2.
$$
 (1)

Here, m is the particle mass, and σ is the particle diameter. This scaling can be understood by dimensional analysis of the rigid granular flow, where the inverse of the shear rate $\dot{\gamma}^{-1}$ is the only time scale in the system. This scaling applies to the normal stress *N* also, which gives

$$
N = m\sigma^{-1}B(\nu)\dot{\gamma}^2.
$$
 (2)

In slope flow under gravity, the force balance gives *S*/*N* =tan θ . Thus we finally have

$$
\frac{S}{N} = \frac{A(\nu)}{B(\nu)} = \tan \theta,\tag{3}
$$

i.e., the packing fraction ν is determined by the inclination angle θ .

This dimensional analysis does not hold when time scales other than $\dot{\gamma}^{-1}$ come into the problem, e.g., the time scales of particle deformation $\lceil 8 \rceil$ $\lceil 8 \rceil$ $\lceil 8 \rceil$, but not only the constant density profile but also the Bagnold scaling itself has been found in numerical simulations of dense steady flow down a slope for hard enough particles $[5]$ $[5]$ $[5]$.

In the slope flow simulations, the value of the packing fraction ν has been shown to increase upon decreasing the inclination angle θ , and eventually the flow stops at a finite angle θ_{stop} ; namely, $A(\nu)/B(\nu)$ is a decreasing function of ν in the dense region $[5,6]$ $[5,6]$ $[5,6]$ $[5,6]$. One can interpret the transition at θ_{stop} as the jamming transition [[3](#page-7-2)[,9](#page-7-7)].

A theoretical analysis of the functional form of $A(\nu)/B(\nu)$ has been done by Louge $[10]$ $[10]$ $[10]$ using kinetic theory, but he found the opposite dependence in the dense region, namely, the theory gives increasing packing fraction ν upon increasing inclination angle θ as shown in Fig. [1,](#page-1-0) where the curves from a kinetic theory $[11]$ $[11]$ $[11]$ are shown by symbols connected by dashed lines for the various restitution coefficients *ep*.

Several explanations for this discrepancy have been proposed, such as enduring contacts $[10,12]$ $[10,12]$ $[10,12]$ $[10,12]$, the Burnett order (the second order of the spatial gradients) effect $[13]$ $[13]$ $[13]$, and the particle roughness $[13]$ $[13]$ $[13]$, etc., but the subject is still under debate.

Recently, the present authors $\lceil 6 \rceil$ $\lceil 6 \rceil$ $\lceil 6 \rceil$ have made a detailed *Permanent address. comparison between the simulation results of dense slope

FIG. 1. (Color online) The ratio of the shear stress to the normal stress S/N vs the packing fraction ν from the simulation data $[e_p = 0.70 \, (\blacksquare), 0.92 \, (\blacksquare),$ and 0.98 (O)] and the plot of Eq. ([20](#page-2-0)) from kinetic theory $[e_p=0.70$ (* connected by dashed line), 0.92 (+ connected by dashed line), and 0.98 (\times connected by dashed line)]. For the simulation data, the average normal stress $N = \left(\frac{1}{3}\right)(N_x + N_y + N_z)$ is used.

flow and the kinetic theory by Jenkins and Richman $[14]$ $[14]$ $[14]$. In contrast with the rather good agreement for the stresses, it has been found that the kinetic theory overestimates the energy dissipation rate Γ , and this discrepancy is responsible for the contradictory behavior in the kinetic theory, i.e., $A(\nu)/B(\nu)$ increases with the packing fraction ν .

The authors conjectured that the discrepancy in the energy dissipation rate Γ should be caused by velocity correlations enhanced by the inelastic collisions; the decrease of the relative normal velocity through the inelastic collisions results in reduction of the energy loss per collision. Such an effect has been noticed in granular gas simulations without shear [15](#page-8-5)[,16](#page-8-6), and the velocity correlations have been investigated analytically $\lceil 17, 18 \rceil$ $\lceil 17, 18 \rceil$ $\lceil 17, 18 \rceil$ $\lceil 17, 18 \rceil$ $\lceil 17, 18 \rceil$.

However, the situation is rather complicated under shear, because the shear tends to break the correlations. The spatial velocity correlations in granular flow under shear has not been carefully studied so far $[19]$ $[19]$ $[19]$.

In this paper, we study the velocity correlation in sheared granular flow, focusing on its effects on the energy dissipation rate and the stress. We adopt the simple shear flow of inelastic hard spheres as a model system, in accordance with most of the kinetic theory analysis. Note that enduring contacts are not allowed in the hard sphere model, whose effects are often under debate in soft sphere model simulations of the slope flow $\vert 5,6 \vert$ $\vert 5,6 \vert$ $\vert 5,6 \vert$ $\vert 5,6 \vert$.

This paper is organized as follows. In Sec. II, we briefly summarize the inelastic hard sphere model and the constitutive relations based on kinetic theory. We summarize our simulation method and present the results in Sec. III. The discussion and a summary are given in Sec. IV.

II. INELASTIC HARD SPHERE MODEL AND KINETIC THEORY

The inelastic hard sphere model is one of the simplest and most widely used models of granular materials $[2,20]$ $[2,20]$ $[2,20]$ $[2,20]$. The

particles are infinitely rigid, and they interact through instantaneous two-body collisions. We adopt the simplest collision rule for the monodisperse smooth hard spheres with diameter σ , mass *m*, and a constant normal restitution coefficient e_p in three dimensions as follows. The particle i at the position r_i with the velocity c_i collides with the particle *j* if $|r_i - r_j| = \sigma$ and $(r_i - r_j) \cdot (c_i - c_j) < 0$, and their postcollisional velocities c_i^* and c_j^* are given by

$$
c_i^* = c_i - \frac{1 + e_p}{2} [n_{ij} \cdot (c_i - c_j)] n_{ij},
$$
 (4)

$$
c_j^* = c_j + \frac{1 + e_p}{2} [n_{ij} \cdot (c_i - c_j)] n_{ij},
$$
 (5)

respectively. Here, n_{ij} is a unit vector defined as $n_{ij} = (r_i)$ $-r_j$)/ $|r_i - r_j|$. The collision is elastic when $e_p = 1$, and inelastic when $0 \lt e_p \lt 1$. In the inelastic case, the particles lose kinetic energy every time they collide; thus an external drive is necessary to keep the particles flowing.

We compare the simulation results of the inelastic hard spheres with the constitutive relations obtained from the Chapman-Enskog method $[1]$ $[1]$ $[1]$, which has been developed in the kinetic theory of gases. In this paper, we employ those by Garzó and Dufty $[11]$ $[11]$ $[11]$, who have improved the previous studies $[2,14,21]$ $[2,14,21]$ $[2,14,21]$ $[2,14,21]$ $[2,14,21]$, which are limited to the weakly inelastic case $[(1-e_p) \ll 1]$, to include the case with any value of the restitution constant e_p under the assumption that the state is near a local homogeneous cooling state $\lceil 22 \rceil$ $\lceil 22 \rceil$ $\lceil 22 \rceil$.

In the following, we briefly summarize the kinetic theory to derive the constitutive relations. The hydrodynamic variables are the number density field $n(r, t)$, the velocity field $u(r, t)$, and the granular temperature field $T(r, t)$, defined in terms of the single-particle distribution function $f(\mathbf{r}, \mathbf{c}, t)$ as

$$
n(\mathbf{r},t) = \int f(\mathbf{r},\mathbf{c},t)d\mathbf{c},\qquad(6)
$$

$$
u(r,t) = \frac{1}{n} \int cf(r,c,t)dc,
$$
 (7)

$$
T(r,t) = \frac{m}{3n} \int (c - u)^2 f(r,c,t) dc.
$$
 (8)

The hydrodynamic equations for these variables are given by

$$
\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{u}) = 0,\tag{9}
$$

$$
mn\frac{\partial u}{\partial t} + mnu \cdot \nabla u = -\nabla \vec{\Sigma},\tag{10}
$$

$$
\frac{3}{2}\left(n\frac{\partial T}{\partial t} + n\mathbf{u} \cdot \nabla T\right) = -\nabla \cdot \mathbf{q} - \tilde{\Sigma} \cdot \vec{E} - \Gamma,\tag{11}
$$

where Σ is the stress tensor, *q* is the heat flux, and *E* is the symmetrized velocity gradient tensor: $E_{ij} = \frac{1}{2} (\partial u_j / \partial r_i)$ + $\partial u_i / \partial r_j$). Note that the energy dissipation rate Γ in Eq. ([11](#page-1-1)) appears due to the energy loss through the inelastic collisions, which gives peculiar features to the granular hydrodynamics.

The constitutive relations for $\vec{\Sigma}$, q , and Γ are determined by the single-particle distribution $f(r, c, t)$. Its time evolution depends on the two-particle distribution function $f^{(2)}(r_1, r_2, c_1, c_2, t)$ through the two-particle collision; the *n*-particle distribution function depends on the $(n+1)$ -particle distribution function. This is known as the Bogoliubov-Born-Green-Kirkwood-Yvon (BBGKY) hierarchy $\lceil 23 \rceil$ $\lceil 23 \rceil$ $\lceil 23 \rceil$.

In the Enskog approximation, the two-particle distribution at collision is approximated as

$$
f^{(2)}(\mathbf{r}_1, \mathbf{r}_1 + \sigma \mathbf{n}_{21}, \mathbf{c}_1, \mathbf{c}_2, t) = g_0(\nu) f(\mathbf{r}_1, \mathbf{c}_1, t) f(\mathbf{r}_1 + \sigma \mathbf{n}_{21}, \mathbf{c}_2, t)
$$
\n(12)

to close the BBGKY hierarchy at the single-particle distribu-tion [[1,](#page-7-0)[11](#page-8-1)]. Here, $g_0(\nu)$ is the radial distribution function at distance σ , and depends on the packing fraction $\nu = \frac{1}{6}\pi\sigma^3 n$. The term $g_0(\nu)$ represents the positional correlations, and the actual procedure to determine the functional form of $g_0(\nu)$ is presented in Sec. III B 1. The correlations in the particle velocities are neglected under the molecular chaos assumption.

The constitutive relations for the hydrodynamic equations have been obtained in Ref. $[11]$ $[11]$ $[11]$ by the Chapman-Esnkog method with the approximation (12) (12) (12) up to the Navier-Stokes order (i.e., the first order of the spatial gradients). In a simple steady shear flow with constant *n*, constant *T*, and $u(r)$ $=(\dot{\gamma}z,0,0)$, the nonzero terms are the pressure or the normal stress

$$
N_{\alpha} \equiv \sum_{\alpha,\alpha} = N(\nu, T) = \sigma^{-3} f_1(\nu) T,\tag{13}
$$

the shear stress

$$
S = \sum_{x,z} = \sum_{z,x} = S(\nu, T) = m^{1/2} \sigma^{-2} f_2(\nu) \sqrt{T} \dot{\gamma}, \qquad (14)
$$

and the energy dissipation rate

$$
\Gamma = \Gamma(\nu, T) = m^{-1/2} \sigma^{-4} f_3(\nu) T^{3/2}.
$$
 (15)

The dimensionless functions $f_i(v)$ are listed in Table [I.](#page-2-2)

In the simple shear flow, Eqs. (9) (9) (9) and (10) (10) (10) are automatically satisfied with the constant normal stress *N* and the constant shear stress S . The energy balance equation (11) (11) (11) gives

$$
S\dot{\gamma} - \Gamma = 0,\t(16)
$$

because there is no heat flux q . Equation (16) (16) (16) means that the granular temperature is locally determined by the balance between the viscous heating and the energy dissipation. Equation (16) (16) (16) with Eqs. (14) (14) (14) and (15) (15) (15) gives

$$
T = m\sigma^2 \frac{f_2(\nu)}{f_3(\nu)} \dot{\gamma}^2.
$$
 (17)

Substituting Eq. (17) (17) (17) into Eqs. (13) (13) (13) and (14) (14) (14) , we get

$$
N = m\sigma^{-1} \frac{f_1(\nu) f_2(\nu)}{f_3(\nu)} \dot{\gamma}^2,
$$
 (18)

TABLE I. The dimensionless functions in the constitutive relations from Ref. $[11]$ $[11]$ $[11]$.

$$
f_1(\nu) \frac{6}{\pi} \nu [1+2(1+e_p) \nu g_0(\nu)]
$$
\n
$$
f_2(\nu) \frac{5}{16\sqrt{\pi}} \left[\gamma^* \left(1 + \frac{4}{5} \nu g_0(\nu)(1+e_p) \right) + \frac{3}{5} \gamma^* \right]
$$
\n
$$
f_3(\nu) \frac{72(1-e_p^2)}{\pi^{3/2}} \nu^2 g_0(\nu) \left(1 + \frac{3}{32} c^* (e_p) \right)
$$
\n
$$
\gamma^{k^*} \left(\nu_\eta^* - \frac{1}{2} \zeta^{(0)*} \right)^{-1} \left(1 - \frac{2}{5} (1+e_p)(1-3e_p) \nu g_0(\nu) \right)
$$
\n
$$
\nu_\eta^*
$$
\n
$$
g_0(\nu) \left(1 - \frac{1}{4} (1-e_p)^2 \right) \left(1 - \frac{1}{64} c^* (e_p) \right)
$$
\n
$$
\zeta^{(0)*} \frac{5}{90} (1 - e_p^2) \left(1 + \frac{3}{32} c^* (e_p) \right)
$$
\n
$$
\gamma^*
$$
\n
$$
\frac{128}{5\pi} \nu^2 g_0(\nu) (1+e_p) \left(1 - \frac{1}{32} c^* (e_p) \right)
$$
\n
$$
c^*(e_p) \frac{32(1-e_p)(1-2e_p^2)[81-17e_p+30e_p^2(1-e_p)]^{-1}}{2}
$$

$$
S = m\sigma^{-1} \frac{[f_2(\nu)]^{3/2}}{[f_3(\nu)]^{1/2}} \dot{\gamma}^2,
$$
\n(19)

which are exactly what we anticipated from the Bagnold scaling Eqs. (1) (1) (1) and (2) (2) (2) .

The above derivation of the Bagnold scaling by the kinetic theory gives a definite expression for Eq. (3) (3) (3) ,

$$
\frac{S}{N} = \frac{\sqrt{f_2(\nu)f_3(\nu)}}{f_1(\nu)},
$$
\n(20)

as a function of the packing fraction ν . This is plotted in Fig. [1](#page-1-0) by symbols connected by lines, along with the simulation data. One can see a clear discrepancy between the theory and the simulation especially in the higher-density region. The kinetic theory gives an increasing function S/N of ν , which means that the flow down a steeper slope is denser.

III. SIMULATIONS

In this section, we compare the expressions Eqs. (13) (13) (13) – (15) (15) (15) with the simulation results for simple shear flow of inelastic hard spheres.

A. Simulation setup

The simulation is done under the constant volume condition with a uniform shear in a rectangular box of the size $L_x \times L_y \times L_z$. The shear is applied by the Lees-Edwards shearing periodic boundary conditions in the *z* direction $[24]$ $[24]$ $[24]$; periodic boundary conditions are employed in the *x* and *y* directions. We employ the event driven method, using the fast algorithm developed by Isobe $\lceil 25 \rceil$ $\lceil 25 \rceil$ $\lceil 25 \rceil$.

A steady shear flow with the mean velocity $u(r)$ $= (\dot{\gamma}_z, 0, 0)$ is prepared as follows. First, a random configuration is prepared by the compressing procedure proposed by Lubachevsky and Stillinger $[26]$ $[26]$ $[26]$ in an elastic system without

 $=$

shear under the periodic boundary conditions. Second, the initial shear flow is constructed from the above random configuration by giving the initial mean velocity $u(r)$ $= (\dot{\gamma}_z, 0, 0)$ and setting the initial temperature $T \approx 100 m \sigma^2 \dot{\gamma}^2$. Last, the steady shear flow of the inelastic system is obtained by relaxing the initial flow under the Lees-Edwards shearing periodic boundary condition $[27]$ $[27]$ $[27]$.

With the present parameters and system size, the final steady state is simple shear flow with uniform packing fraction $\nu = \nu_0$ and mean velocity $\mathbf{u} = (\dot{\gamma}_z, 0, 0)$ [[28](#page-8-18)]. All the following data are taken in the steady state, and averaged over space and time (typically over 10 000 collisions per particle) unless otherwise noted.

In the following, all the quantities are given in dimensionless form with the unit mass m , the unit length σ , and the unit time $\dot{\gamma}^{-1}$. Most of the data are from the simulations with system size $L_x = 20$, $L_y = 10$, and $L_z = 40$. Several simulations has been done with $L_x = L_y = L_z = 40$ to check the system size effect. We measure the temperature *T*, the normal stress *N*, the shear stress *S*, and the energy dissipation rate Γ for various values of the packing fraction ν . These are compared with Eqs. (13) (13) (13) – (15) (15) (15) from the kinetic theory.

B. Simulation results

1. The radial distribution function

For the constitutive relations of Table [I,](#page-2-2) we need to know the radial distribution function at the particle diameter, $g_0(\nu)$, as a function of the packing fraction. For elastic hard spheres $(e_p=1)$ in equilibrium, the well-known expression of $g_0(\nu)$ is the Carnahan-Starling formula $[24]$ $[24]$ $[24]$

$$
g_{0,\text{CS}}(\nu) = \frac{1 - \nu/2}{(1 - \nu)^3} \tag{21}
$$

for $0 < \nu < \nu_f$, where ν_f is the freezing packing fraction and $\nu_f \approx 0.49$ [[29](#page-8-19)]. Torquato [29] proposed a formula that includes the higher packing fraction up to the random close packing fraction $\nu_c \approx 0.64$ as

$$
g_{0,T}(\nu) = \begin{cases} g_{0,CS}(\nu) & \text{for } 0 < \nu < \nu_f, \\ g_{0,CS}(\nu_f)(\nu_f - \nu_c)/(\nu - \nu_c) & \text{for } \nu_f < \nu < \nu_c. \end{cases}
$$
(22)

As for inelastic hard spheres under shear, a generally accepted form of $g_0(\nu)$ does not exist, but it has been found in several simulations that $g_0(\nu)$ is larger for stronger inelastic-ity [[15](#page-8-5)[,30](#page-8-20)]. Figure [2](#page-3-0)(a) shows the radial distribution $g(r)$ averaged over all directions obtained from our shear flow simulation with the packing fraction $\nu = 0.58$ for various values of e_p . The spatial mesh to measure $g(r)$ was taken as 0.001, and the peak values of $g(r)$ around $r=1$ (at the distance of the particle diameter) are marked by symbols for e_p = 0.98 and 0.92. We can see that the peak value strongly increases for smaller e_p , and can be much larger than the value from Eq. (22) (22) (22) $\left[g_{0,T}(0.58),\right]$ shown by an arrow. It is quite difficult to evaluate the precise value of $g_0(\nu)$ from this direct measurement of $g(r)$ because of the strong increase of $g(r)$ in the limit of $r \rightarrow +1$.

FIG. 2. (Color online) (a) Radial distribution functions for ν $= 0.58$ with $e_p = 0.7$ (dashed line), 0.92 (dotted line), and 0.98 (solid line). The peak values of contact are 136 (out of range), 30.6 $\textcircled{\bullet}$), and 17.4 (\circ) for $e_p = 0.70$, 0.92, and 0.98, respectively. The theoretical value at contact, $g_{0,T}(0.58)$, is shown by an arrow. (b) Plot of $g_{0,\text{m}}(\nu)$ vs the packing fraction ν for e_p =0.7 (\blacksquare), 0.92 (\blacklozenge), and 0.98 (O). $g_{0,T}(\nu)$ is shown by a solid line.

The way we determine $g_0(\nu)$ from the simulation is through the expression of the collision frequency ω_0 [[31,](#page-8-21)[32](#page-8-22)] from the kinetic theory $[11,33]$ $[11,33]$ $[11,33]$ $[11,33]$,

$$
\omega_0(\nu, T) = 24g_0(\nu)\sqrt{T}\nu\pi^{-1/2}\left(1 - \frac{1}{32}c^*(e_p)\right),\tag{23}
$$

where $c^*(e_p)$ is given in Table [I.](#page-2-2) By measuring ω_0 and *T* for each ν from the simulation, we can evaluate

$$
g_{0,m}(\nu;T,\omega_0) \equiv \frac{\omega_0 \sqrt{\pi}}{24[1-c^*(e_p)/32]\sqrt{T}\nu}.
$$
 (24)

 $g_{0,m}(\nu; T, \omega_0)$ is plotted versus ν for various values of e_p in Fig. [2](#page-3-0)(b), where $g_{0,T}(\nu)$ in Eq. ([22](#page-3-1)) is shown by a solid line for reference. $g_{0,m}(\nu; T, \omega_0)$ shows a stronger increase upon increasing the packing fraction ν as e_p gets smaller; by comparing it with Fig. $2(a)$ $2(a)$, we see that this indirect estimate gives a reasonable e_p dependence of $g_0(\nu)$. In the following, we use $g_{0,m}(\nu; T, \omega_0)$ as $g_0(\nu)$ in Table [I](#page-2-2) unless otherwise noted.

2. The energy dissipation rate and the normal stress as functions of the packing fraction

In Fig. [3,](#page-4-0) the energy dissipation rate Γ [Fig. [3](#page-4-0)(a)] and the normal stress N [Fig. $3(b)$ $3(b)$] are shown for various values of the packing fraction ν and the restitution coefficient e_p . For the normal stress, we find in the simulation that N_a depends on the direction α , but the differences among them are at most 10% in the plotted region and are not significant compared to the difference from the kinetic theory that we will study in the following. Thus, here we plot the average *N* $\equiv (N_x + N_y + N_z)/3.$

The values from kinetic theory are shown in Figs. $3(a)$ $3(a)$ and $3(b)$ $3(b)$ by symbols connected by dashed lines. We see in Fig. $3(a)$ $3(a)$ that the energy dissipation rate is overestimated by the theory in the dense region, and the disagreement is larger for smaller e_p . The normal stress in Fig. $3(b)$ $3(b)$ also shows a similar tendency, although the relative disagreements are smaller than those in the energy dissipation rate Γ .

FIG. 3. (Color online) Energy dissipation rate Γ (a) and the normal stress *N* (b). The simulation data for $e_p = 0.98$ (O), 0.92 (\bullet), and 0.70 (\blacksquare) are compared with the values from the kinetic theory $[\Gamma(\nu, T)$ and $N(\nu, T)]$ shown by symbols connected by dashed lines for $e_p = 0.98$ (\times), 0.92 (+), and 0.70 (*). $\Gamma(\nu, T_{\text{coll}})$ and $N(\nu, T_{\text{coll}})$ with $g_{0,m}(\nu; T_{\text{coll}}, \omega_0)$ are denoted by the solid lines, which agree with the simulation data (see text).

3. The precollisional velocity correlation effects and the collisional temperature

a. The energy dissipation. We first focus on the discrepancy in the energy dissipation rate Γ . From the collision rule Eq. (5) (5) (5) , the energy dissipated per collision is given by

$$
\Delta E_{ij} = (1 - e_p^2) \frac{1}{4} c_{n,ij}^2,\tag{25}
$$

where $c_{n,ij} \equiv [(c_i - c_j) \cdot n_{ij}]$ is the relative normal velocity of colliding particles just before the collision. Thus, Γ is given by

$$
\Gamma = \langle \Delta E_{ij} \rangle_{\text{coll}} \frac{1}{2} n \omega_0 = (1 - e_p^2) \frac{1}{4} \langle c_{n,ij}^2 \rangle_{\text{coll}} \frac{1}{2} n \omega_0.
$$
 (26)

Here, $\langle A \rangle_{\text{coll}}$ denotes the average of a quantity *A* over all collisions; if the value of *A* is A_k at the *k*th collision, $\langle A \rangle_{\text{coll}}$ $\equiv \sum_{k=1}^{N_{\text{coll}}} A_k / N_{\text{coll}}$, where N_{coll} is the total number of collisions. Note that Eq. (26) (26) (26) is the exact expression for Γ .

On the other hand, expression (15) (15) (15) from the kinetic theory with Eq. (24) (24) (24) gives

FIG. 4. (Color online) Temperature T and collisional temperature T_{coll} vs the packing fraction ν . *T* and T_{coll} are denoted by \blacksquare and * with the dashed lines for $e_p = 0.7$, respectively, and by \bullet and $+$ with the dashed lines for $e_p = 0.92$. The inset shows *T* and T_{coll} for e_p =0.98 represented by \circ and \times with the dashed lines.

$$
\Gamma(\nu, T) = (1 - e_p^2) T \left(\frac{1 + 3c^*(e_p)/32}{1 - c^*(e_p)/32} \right) \frac{1}{2} n \omega_0(\nu, T). \quad (27)
$$

To interpret this expression, let us consider the random collision of particles whose velocity fluctuation is given by the Maxwellian. In this case, $\frac{1}{4} \langle c_{n,ij}^2 \rangle_{\text{coll}} = T$; then Eq. ([26](#page-4-1)) gives

$$
\Gamma = (1 - e_p^2) T_{\frac{1}{2}}^{\frac{1}{2}} n \omega_0.
$$
 (28)

The difference between this and Eq. (27) (27) (27) comes from the deviation of the velocity distribution from the Maxwellian, but the difference is found to be small in the parameter region studied in the present paper. Therefore, from the comparison of the exact expression (26) (26) (26) with the kinetic theory expression Eq. (27) (27) (27) , we conclude that the deviation found in Fig. [3](#page-4-0)(a) comes from the fact that $\frac{1}{4} \langle c_{n,ij}^2 \rangle_{\text{coll}} \ll T$.

b. The collisional temperature. To confirm this idea, we define the "collisional temperature" $T_{\text{coll}} = \langle c_{n,ij}^2 \rangle_{\text{coll}}/4$. Figure [4](#page-4-3) shows T_{coll} and *T* as functions of ν . One can see that T_{coll} is substantially smaller than T for $\nu > 0.5$ as is concluded above.

To demonstrate that the discrepancy is actually resolved by T_{coll} , we plot $\Gamma(\nu, T)$ of Eq. ([15](#page-2-5)) with $g_{0,m}(\nu, T, \omega_0)$ of ([24](#page-3-2)) in $f_3(v)$ $f_3(v)$ $f_3(v)$ replacing *T* by T_{coll} [the solid lines in Fig. 3(b)]; this is equivalent to replacing *T* in Eq. ([27](#page-4-2)) with T_{coll} and using the measured value of the collision frequency for ω_0 . The agreement is quite good.

c. Normal stress. Now we consider the effect of T_{coll} $\leq T$ on the normal stress *N*. The value of $c_{n,ij}$ should also play an important role in the collisional component of the normal stress N_{coll} , because $c_{n,ij}$ directly determines the momentum transfer from the particle *i* to the particle *j* through a collision: $\Delta p_j = (c_j^* - c_j) = [(1 + e_p)/2]c_{n,ij}n_{ij} = -\Delta p_i$. Thus, we expect that N_{coll} is approximately proportional to $\langle |\Delta p| \rangle n \omega_0$ $\propto \sqrt{T_{\text{coll}}n\omega_0}$. In addition, the collisional part *N*_{coll} is dominant

FIG. 5. (Color online) Distribution of relative normal velocity just before the collision c_n . $P_{all}(c_n)$ and $P_{interval}(c_n)$ are compared along with $P_{\text{rand}}(c_n; T_{\text{coll}})$ and $P_{\text{rand}}(c_n; T)$ for $e_p = 0.92$ with ν $= 0.40$ (a) and 0.58 (b). We see that the difference between $P_{all}(c_n)$ (O) and $P_{\text{interval}}(c_n)$ (\square) is small for $\nu=0.40$, but for $\nu=0.58$, the $P_{\text{all}}(c_n)$ has sharper distribution. The solid lines show $P_{\text{rand}}(c_n; T_{\text{coll}})$ and the dashed lines show $P_{\text{rand}}(c_n; T)$. See text for details.

in the dense region.

In Fig. $3(b)$ $3(b)$, $N(\nu, T_{\text{coll}})$ of Eq. ([13](#page-2-7)) is plotted by the solid lines, where $g_{0,m}(\nu; T_{coll}, \omega_0)$ is used as $g_0(\nu)$. We see that the solid lines show reasonably good agreement with the data for the whole density region.

4. Origin of the precollisional velocity correlation

One of the possible origins of the precollisional velocity correlation that makes $T_{\text{coll}} < T$ is the inelasticity, which makes the relative normal velocity smaller upon collision. In this section, we examine how the precollisional velocity correlation develops in the shear flow.

It is expected that the correlation grows when particles collide with the same colliding partners inelastically many times within a short period of time. Under shear, however, this correlation will be lost when they are forced to pass each other and collide with new partners. The typical time scale for a pair of particles passing each other is the unit time, i.e., $\dot{\gamma}^{-1}$. This argument explains the smaller T_{coll} in the denser region, because particles collide more frequently with the same partners before they move far apart $\lceil 34 \rceil$ $\lceil 34 \rceil$ $\lceil 34 \rceil$.

FIG. 6. (Color online) Precollisional relative normal velocity distributions $P_{\text{all}}(c_n)$ (\bigcirc) and $P_{\text{interval}}(c_n)$ (\bigcirc) compared along with $P_{\text{rand}}(c_n; T_{\text{coll}})$ (solid line) and $P_{\text{rand}}(c_n; T)$ (dashed line), respectively, for $e_p = 0.70$ with $\nu = 0.58$.

This argument shows that the collision does not have memories of the previous collisions earlier than the unit time $\dot{\gamma}^{-1}$. To confirm this, we compare the following two distributions of the precollisional velocity: (i) $P_{all}(c_n)$, which is the distribution of $c_{n,ij}$ for all collisions between all pairs of particles, and (ii) $P_{\text{interval}}(c_n)$, which is the distribution of $c_{n,ij}$ of the collisions whose colliding pairs of particles did not collide with each other during the last unit time $\dot{\gamma}^{-1}$. If the velocity correlation mainly comes from the multiple collision with the same partners within the unit time scale, then $P_{\text{interval}}(c_n)$ should have the width determined not by T_{coll} but by the average temperature *T*.

The results are shown in Fig. [5](#page-5-0) for $e_p = 0.92$ with ν $= 0.40$ [Fig. [5](#page-5-0)(a)] and 0.58 [Fig. 5(b)], where $P_{all}(c_n)$ is denoted by \bigcirc and $P_{\text{interval}}(c_n)$ is denoted by \Box . We see that $P_{\text{interval}}(c_n)$ is wider than $P_{\text{all}}(c_n)$ for the denser case [Fig. $5(b)$ $5(b)$].

If the particles with the Maxwellian velocity distribution with temperature \tilde{T} collide among themselves randomly, the distribution of c_n is given by

$$
P_{\text{rand}}(c_n; \widetilde{T}) = \frac{c_n}{2\widetilde{T}} \exp\left(-\frac{c_n^2}{4\widetilde{T}}\right). \tag{29}
$$

In Fig. [5,](#page-5-0) $P_{rand}(c_n; T)$ and $P_{rand}(c_n; T_{coll})$ are shown by the dashed and solid lines, respectively. They are indistinguishable for ν =0.40 in Fig. [5](#page-5-0)(a), but show clear difference for ν =0.[5](#page-5-0)8 in Fig. 5(b). We find that $P_{\text{rand}}(c_n; T)$ fits $P_{\text{interval}}(c_n)$, and $P_{\text{rand}}(c_n; T_{\text{coll}})$ fits $P_{\text{all}}(c_n)$, which further confirms that colliding partners are correlated in the way characterized by T_{coll} .

For smaller *ep*, the shape of the distributions deviates from Eq. (29) (29) (29) based on the random collision. Figure [6](#page-5-2) shows $P_{\text{all}}(c_n)$ and $P_{\text{interval}}(c_n)$ compared along with $P_{\text{rand}}(c_n; T_{\text{coll}})$ and $P_{\text{rand}}(c_n; T)$, respectively, for $e_p = 0.70$ with $\nu = 0.58$. $P_{\text{all}}(c_n)$ has sharper distribution than $P_{\text{interval}}(c_n)$, but neither of them fit well with $P_{\text{rand}}(c_n; T_{\text{coll}})$ nor $P_{\text{rand}}(c_n; T)$. This suggests a stronger correlation than the case of $e_p = 0.92$.

FIG. 7. (Color online) Spatial velocity correlation functions for e_p =0.92 and ν =0.55. (a) The correlations in the *x* direction $C_{\alpha,\beta}^{(x)}$ $X(R)$. One sees that the longitudinal component $C_{x,x}^{(x)}(R)$ (solid line) has a larger amplitude than the others (shown by dashed lines). (b) The longitudinal velocity correlations $C_{\alpha,\alpha}^{\alpha}(R)$ for $\alpha = x$ (solid line), y (dashed line), and z (dotted line).

5. The spatial correlation in the velocity fluctuation

To understand the velocity correlations in more detail, we study the spatial velocity correlation function defined as

$$
C_{\alpha,\beta}^{(\gamma)}(R) = \frac{\left\langle \sum_{i,j} \left\{ \tilde{c}_{\alpha,i} \tilde{c}_{\beta,j} \Delta[R - |(r_i - r_j) \cdot e_{\gamma}|] \right\} \right\rangle}{\left\langle \sum_{i,j} \left\{ \Delta[R - |(r_i - r_j) \cdot e_{\gamma}|] \right\} \right\rangle}, \quad (30)
$$

where α , β , and γ are *x*, *y*, or *z*, $\tilde{c}_{\alpha,i} \equiv (c_{\alpha,i} - u_{\alpha}), \langle \cdots \rangle$ denotes the time average, and e_{γ} represents the unit vector in the γ direction. $\Delta (R - |r \cdot e_y|)$ is 1 when $|R - |r \cdot e_y|| < 0.05$ and $|r \cdot e_{\gamma'}|$ < 0.1 for $\gamma' \neq \gamma$; otherwise it is zero. We calculated $C_{\alpha,\beta}^{(\gamma)}(R)$ for the system with e_p =0.92, for both a small system with $L_x=20$, $L_y=10$, $L_z=40$ and a large system with L_x $= 40, L_y = 40, L_z = 40$. We find that the correlation extends over the whole system in the case of the small system, but it goes to zero for the large system. In the following, we present the spatial correlation measured in the large system, but we confirmed that the hydrodynamic quantities presented in the previous sections did not show any differences.

In Fig. $7(a)$ $7(a)$, the various components of the correlation in the *x* direction $C_{\alpha,\beta}^{(x)}(R)$ are shown. We find that the longitu-

FIG. 8. (Color online) Shear stress S vs packing fraction ν . The simulation data are plotted for $e_p = 0.98$ (O), 0.92 (\bullet), and 0.70 (\blacksquare) . The kinetic theory constitutive relations *S*(ν , *T*) are shown by the symbols connected by dashed lines (\times for e_p =0.98, + for 0.92, and $*$ for $e_p = 0.70$).

dinal correlation in the *x* direction, $C_{x,x}^{(x)}(R)$, has larger amplitude than the other components; this tendency is also found in the y and z directions (data not shown). The longitudinal correlation at the particle diameter distance $(R=1)$ is positive, which is consistent with the fact that $T_{\text{coll}} < T$. It is evident that the correlation shows an oscillation whose wavelength is of the order of the particle diameter, which will be discussed in Sec. IV.

The longitudinal components in the *x*, *y*, and *z* directions are shown in Fig. $7(b)$ $7(b)$. All of them show oscillations in the particle diameter scale. We also found that the longitudinal correlation shows a larger amplitude for a smaller restitution coefficient e_p and/or larger packing fraction ν (data not shown).

6. The packing fraction dependence of the shear stress

We find that the shear stress *S* shows more complicated packing fraction ν dependence than those of the energy dissipation rate Γ and the normal stress *N*. In Fig. [8,](#page-6-1) the simulation data of the shear stress *S* are denoted by symbols, and $S(\nu, T)$ from the kinetic theory [Eq. ([14](#page-2-4)) with Table [I](#page-2-2)] are denoted by symbols with dashed lines. We find that, for e_n $= 0.98$, the shear stress is underestimated by the theory, while for $e_p = 0.92$ and 0.70, the shear stress is overestimated.

Actually, in the case of the elastic $(e_p = 1)$ hard sphere system, the Enskog theory is known to underestimate the shear viscosity in the dense region $[23,35]$ $[23,35]$ $[23,35]$ $[23,35]$, and this tendency is seen in the result for $e_p = 0.98$. The result for $e_p = 0.70$ shows that the inelasticity reduces the shear stress to a value smaller than the one expected from the kinetic theory, but we do not understand the reason for this reduction yet. The rather good agreement in between for the case of $e_p = 0.92$ seems to be accidental.

IV. DISCUSSION AND SUMMARY

A. The shear stress and the anisotropic correlation

In contrast to the energy dissipation rate Γ and the normal stress *N*, the discrepancy in the shear stress *S* cannot be understood just by the precollisional velocity distribution averaged over all directions, but the anisotropy of the precollisional correlations in both the velocity and the position should be important in the shear stress. These anisotropies are not taken into account in the kinetic theory employed in the present analysis. In fact, for a soft sphere system in twodimensional, sheared flow, it has been found that the contact force distribution strongly depends on direction $[36]$ $[36]$ $[36]$. Our preliminary results also show a similar direction dependence in the collisional momentum transfer per unit time. A detailed analysis is left for future studies.

B. The packing fraction dependence of the ratio of the shear stress to the normal stress

As we have seen in Fig. [1,](#page-1-0) *S*/*N* in the simulation is a decreasing function of the packing fraction ν for larger packing fraction ν and/or smaller restitution coefficient e_p , while in Eq. ([20](#page-2-0)), from the kinetic theory, *S/N* always increases with ν .

Kumaran argued that particle roughness is necessary for S/N to have a decreasing part upon increasing ν in the dense region $\left[13\right]$ $\left[13\right]$ $\left[13\right]$. However, even for smooth particles, the present simulations show that *S*/*N* has a decreasing part in the dense region for the inelastic hard sphere system, although the particle roughness may well amplify the decreasing part of *S*/*N*.

The present authors have suggested $\lceil 6 \rceil$ $\lceil 6 \rceil$ $\lceil 6 \rceil$ that kinetic theory gives increasing S/N with ν even for the denser region because $f_3(v)$ in the energy dissipation Γ of Eq. ([15](#page-2-5)) increases too sharply for larger ν . In this paper in Sec. III B 3, we showed that the sharp increase in Γ can be weakened by using T_{coll} instead of *T*. In the present treatment, however, it is not possible to extract the ν dependence out of $\Gamma(\nu, T_{\text{coll}})$ and to compare it directly with $f_3(v)$ because *T* and T_{coll} are determined by ν and $\dot{\gamma}$ in the steady state simulations; therefore, the quantity that corresponds to $f_3(v)$ in Eq. ([15](#page-2-5)) cannot be defined from the simulation data.

Finally, let us comment on the fact that *S*/*N* does increase with ν in a certain parameter range in our simple shear flow simulation, in contrast to the fact that increasing ν upon increasing S/N =tan θ has never been observed in granular flow down a slope. This suggests that the steady flow in this parameter region is unstable in the slope flow configuration. It would be interesting to study the relation between the stability of the flow and the ν dependence of S/N .

C. Oscillation in the spatial velocity correlation

As shown in Fig. [7,](#page-6-0) the spatial velocity correlation is found to oscillate in the scale of the particle diameter. Although we have not yet understood the origin of this oscillation, it is plausible that the oscillation comes from the coupling between the density correlation and the velocity correlation. Analysis of the sheared Langevin system suggests that the spatial velocity correlation is related to the radial distribution function $\left[37\right]$ $\left[37\right]$ $\left[37\right]$, which oscillates in the particle diameter scale. It is likely that similar coupling also exists in granular shear flow.

D. Summary

We have simulated the simple shear flow of a smooth inelastic hard sphere system by molecular dynamics simulations. We have found that the energy dissipation rate Γ and the normal stress *N* are smaller than those expected from the kinetic theory. We have showed that the relative precollisional normal velocity of colliding pairs of particles, $c_{n,ii}$, is smaller than the one expected from random collisions, and this reduces Γ and *N*. By examining the distributions of $c_{n,ii}$ for all collisions $[P_{\text{all}}(c_n)]$ and for only the first collisions of the new pairs during the last period of time $\dot{\gamma}^{-1}$ [$P_{\text{interval}}(c_n)$], we have concluded that the reduction of the relative velocity is caused by multiple inelastic collisions during the time period $\dot{\gamma}^{-1}$.

To understand the velocity correlation in more detail, we have studied the spatial velocity correlation. It has been found that the longitudinal components of the correlations have larger amplitude with the oscillation in the scale of the particle diameter.

The shear stress *S* has been found to be overestimated for smaller e_p , but underestimated for larger e_p by the kinetic theory.

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